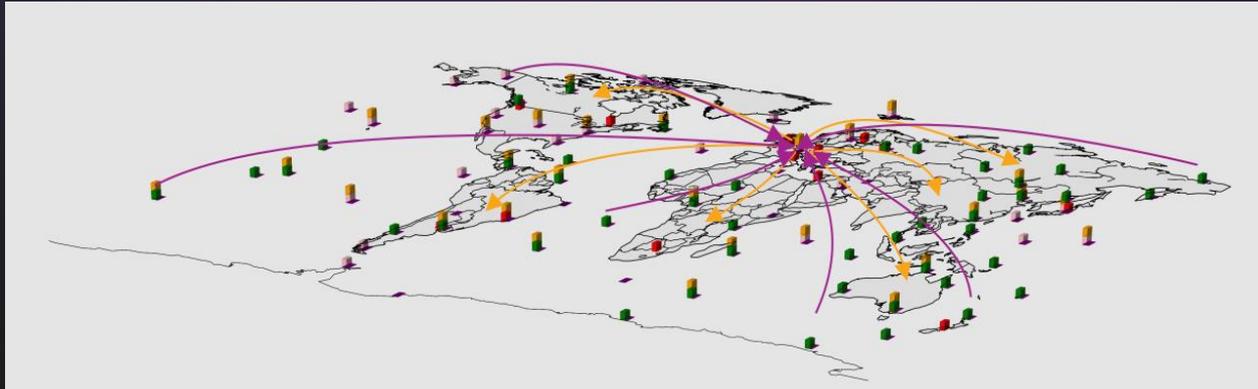




The IMS radionuclide network- a unique machine not yet fully exploited

Anders Ringbom
Swedish Defence Research Agency (FOI)



The IMS RN network is unique

- The only global detection system for atmospheric radioactivity
- Raw data as well as analyzed data is globally distributed.
- The sensors have higher time resolution than in other networks
- Many particulate stations are automatic
- It contains a network of noble gas stations (all automatic)



But..

- The system was designed 30 years ago
- Releases from UG tests shown to be smaller than expected
- The knowledge of the global background was limited at the time, in particular for xenon
- The design did not take the entire verification process into account
- The number of stations, as well as their placement, was the result of negotiations



Outline

- The verification mission
- The RN signal
- The measurement system
- Observations and experience the last 20 years
- Future development

Special thanks to:

Kurt Ungar
Harry Miley
Johan Kastlander
Hakim Gheddou

However, all views expressed are my own





The verification mission

To detect nuclear explosions by performing:

Detection → Location → Categorization

The analysis should be *consistent* and *coherent*

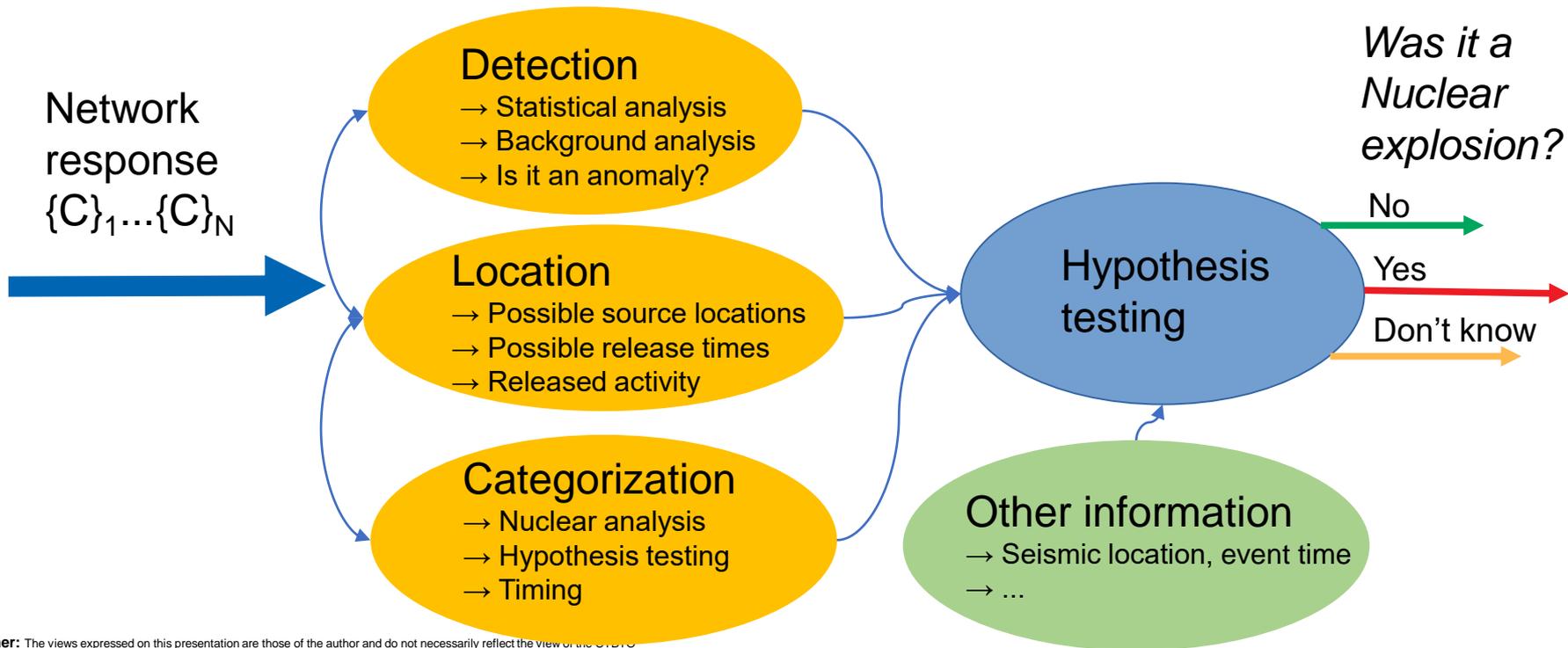


The verification mission

Prove with enough confidence that a nuclear test occurred,
applying a set of pre-defined criteria

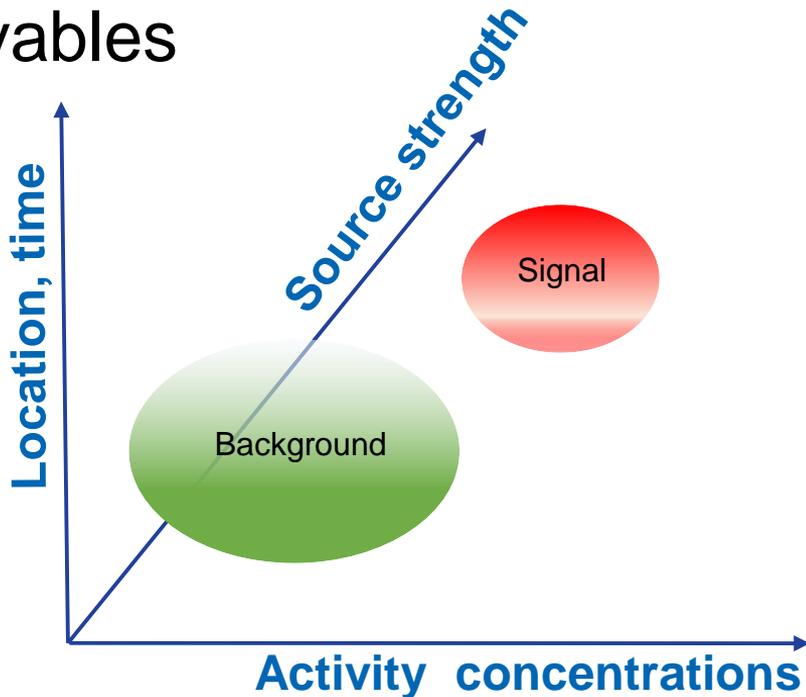
~~Find a nuclear explosion scenario that can explain the observations~~

The RN verification process

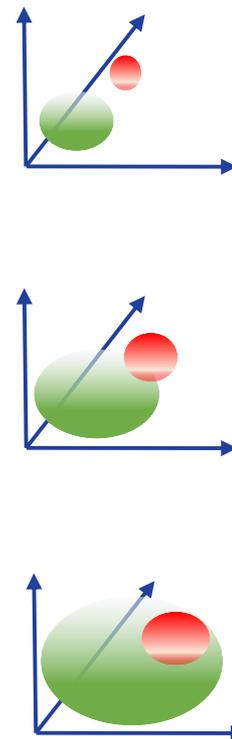


Disclaimer: The views expressed on this presentation are those of the author and do not necessarily reflect the view of the CTBTO

Signal vs background in space of observables

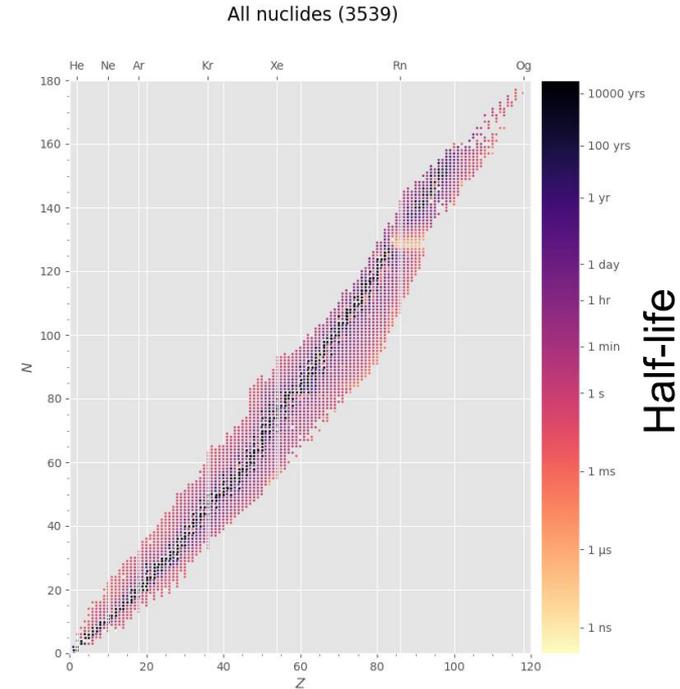


Improving the verification system



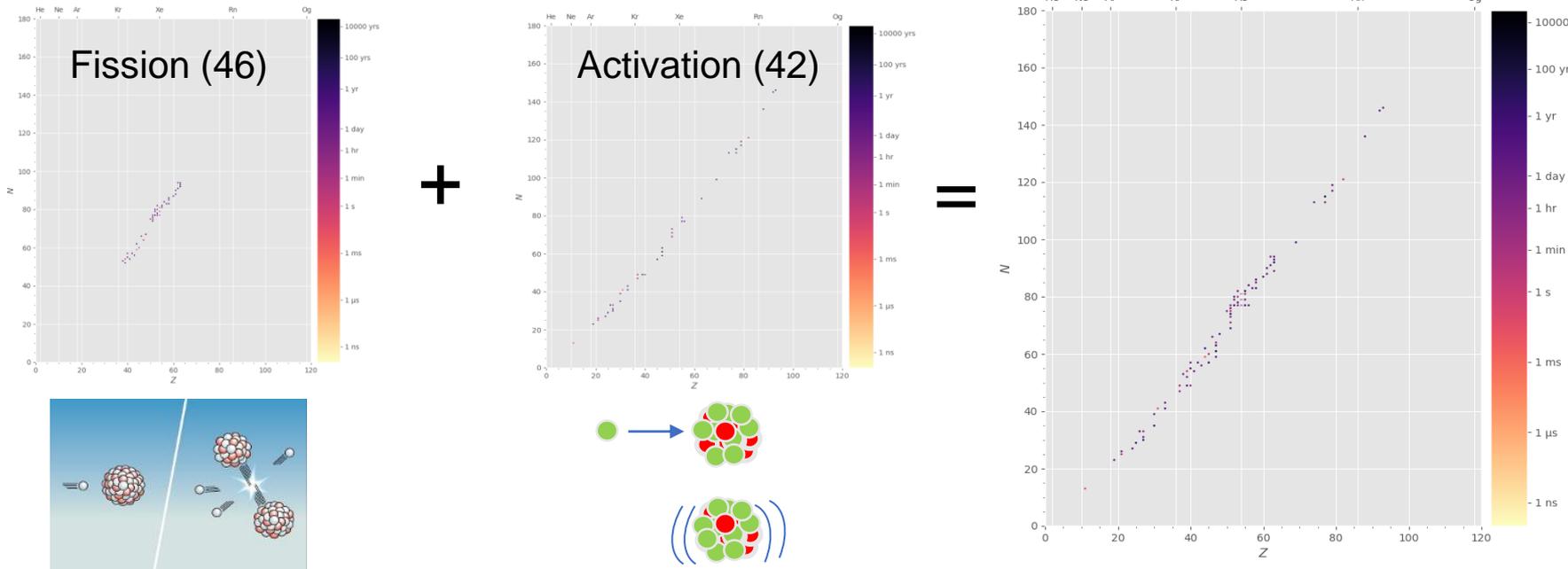
The RN – signal. What is relevant for CTBT?

- Radionuclides formed in a nuclear explosion in high enough quantities
- Half-lives long enough to allow them to be transported and measured
- Detectable gamma radiation
- List should be agreed among member states



88 CTBT-relevant radionuclides

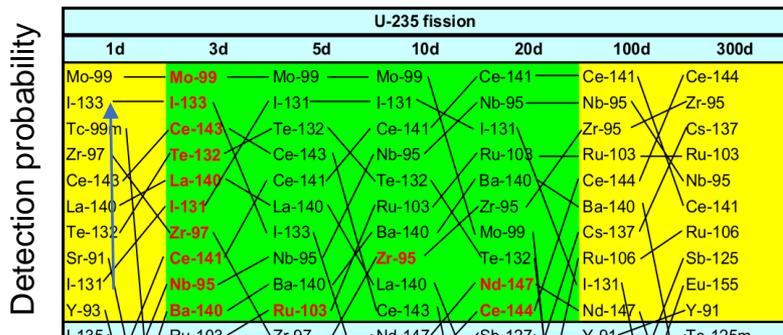
CTBT relevant nuclides (88)



Disclaimer: The views expressed on this presentation are those of the author and do not necessarily reflect the view of the CTBTO

The signal – 88 CTBT-relevant radionuclides

- But some nuclides are more probable to detect than others
Depends on for example production yield, branching ratios, and volatility
- The most probable scenario is an underground NE.



* K M Matthews, NRL-report 2005/1

Defines "significant" nuclides

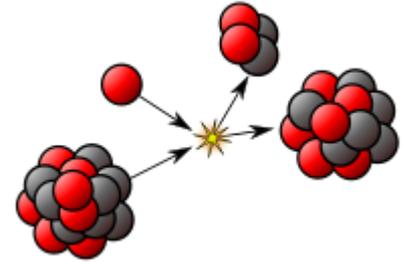
Tabulated from DOE/NV-317

Noble Gas			Volatile above 600° C			Refractory					
Isotope	Number	T _{1/2}	Cum. Yield	Isotope	Number	T _{1/2}	Cum. Yield	Isotope	Number	T _{1/2}	Cum. Yield
^{131m} Xe	16	11.8 d	0.045	¹³¹ I	131	8.03 d	3.22	⁹⁹ Mo	3	65.9 h	5.94
¹³³ Xe	310	5.25 d	6.72	¹³² I	13	2.30 h	4.67	^{99m} Tc	1	6.01 h	5.23
¹³⁵ Xe	169	2.20 d	0.192	¹³³ I	109	20.8 h	6.72	¹³⁹ Ba	6	82.9 m	6.34
¹³⁵ Xe	271	9.14 h	6.60	¹³⁴ I	6	52.5 m	7.64	¹⁴⁰ Ba	19	12.8 d	5.98
				¹³⁵ I	88	6.58 h	6.30	¹⁴⁰ La	17	1.68 d	5.98
				¹³² Te	13	3.20 d	4.66				
				¹³⁷ Cs	10	30.1 y	6.22				
				¹³⁸ Cs	29	32.5 m	6.65				
				¹³⁹ Cs	1	9.27 m	6.32				

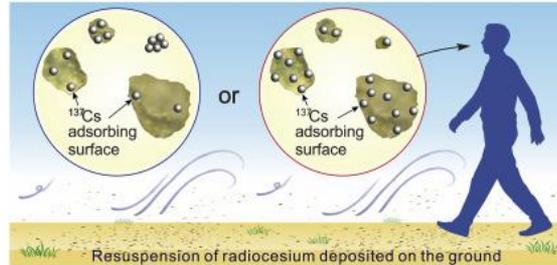
Harry Miley, Paul Eslinger, Ramesh Sarathi, *Impact of environmental background on atmospheric monitoring of nuclear explosions*, talk at WOSMIP remote II, 2021

The radionuclide background

- Background from re-suspended soil: Rn and Th, ^{40}K , ^{238}U
- Cosmogenic: ^7Be (spallation of N,O), $^{22,24}\text{Na}$ (spallation of ^{40}Ar)
- Anthropogenic: IPFs, NPPs, research, industry, accidents, historical tests
- Background from the detector itself



Forsmark NPP, Sweden



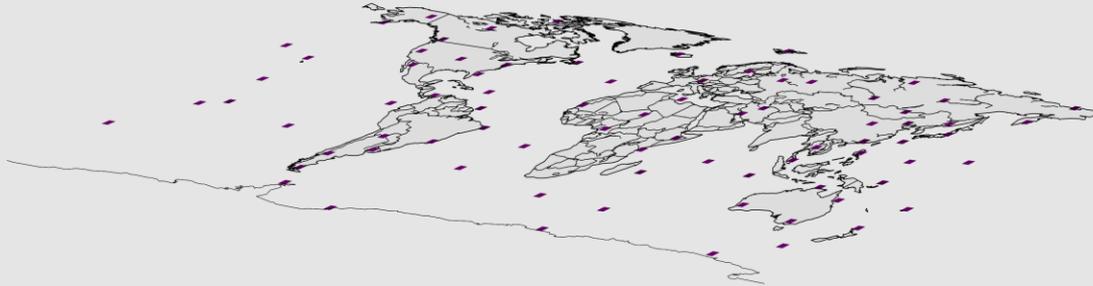
N. Kaneyasu, *et al.*, <https://doi.org/10.1016/j.jenvrad.2017.03.001>

Disclaimer: The views expressed on this presentation are those of the author and do not necessarily reflect the view of the CTBTO



The machine

79 (80) globally distributed sites

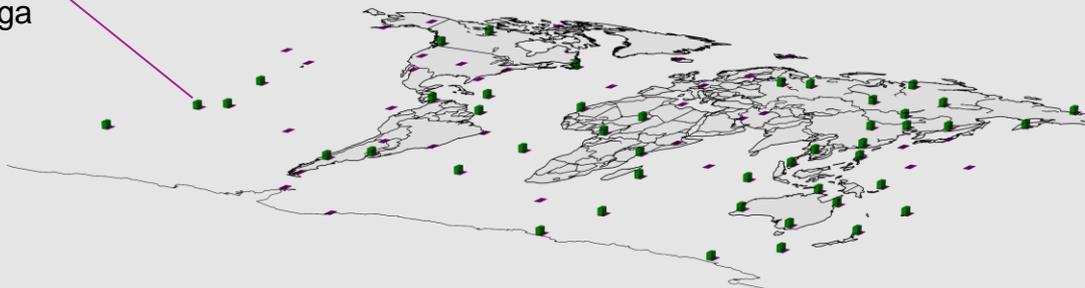


46 certified manual particulate stations



RN23, Raratonga

Collection time:	24 h
Decay time:	24 h
Measurement time:	24 h
Air flow:	> 500 m ³ /h
Sensitivity	< 10-30 μBq/m ³ for ¹⁴⁰ Ba



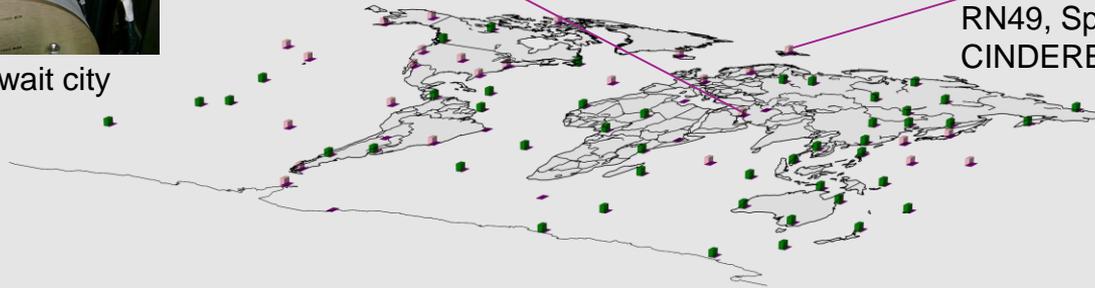


RN40, Kuwait city
RASA

27 certified automatic particulate stations



RN49, Spitzbergen
CINDERELLA



26 certified noble gas stations

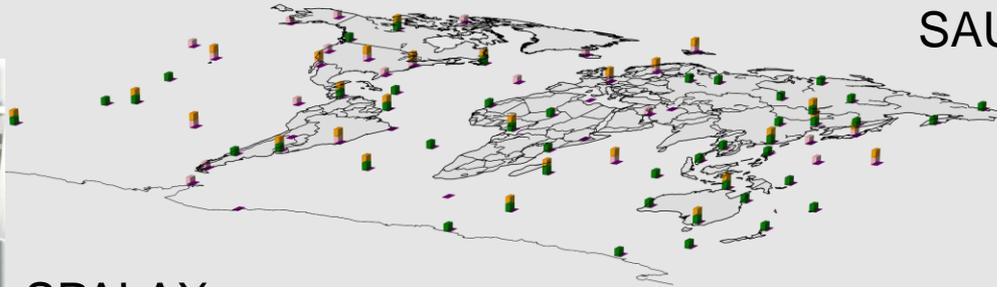
Collection time:	12 or 24 h
Decay time:	~6 h
Measurement time:	11 or 24 h
Air flow:	> 1 m ³ /h
MDC ¹³³ Xe:	~0.3 mBq/m ³



SAUNA II



SPALAX

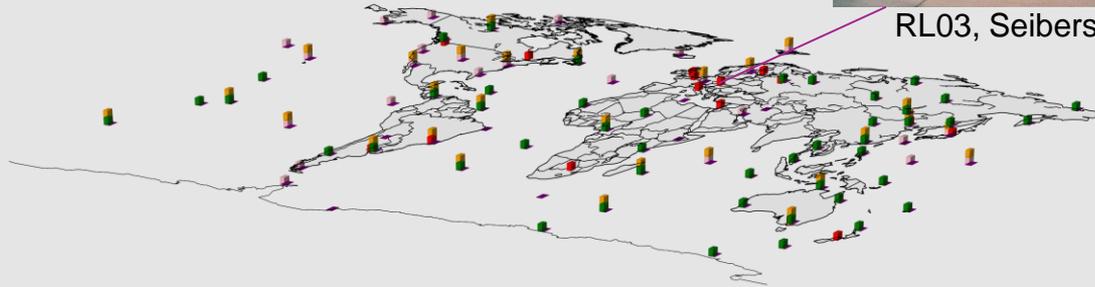


16 radionuclide laboratories

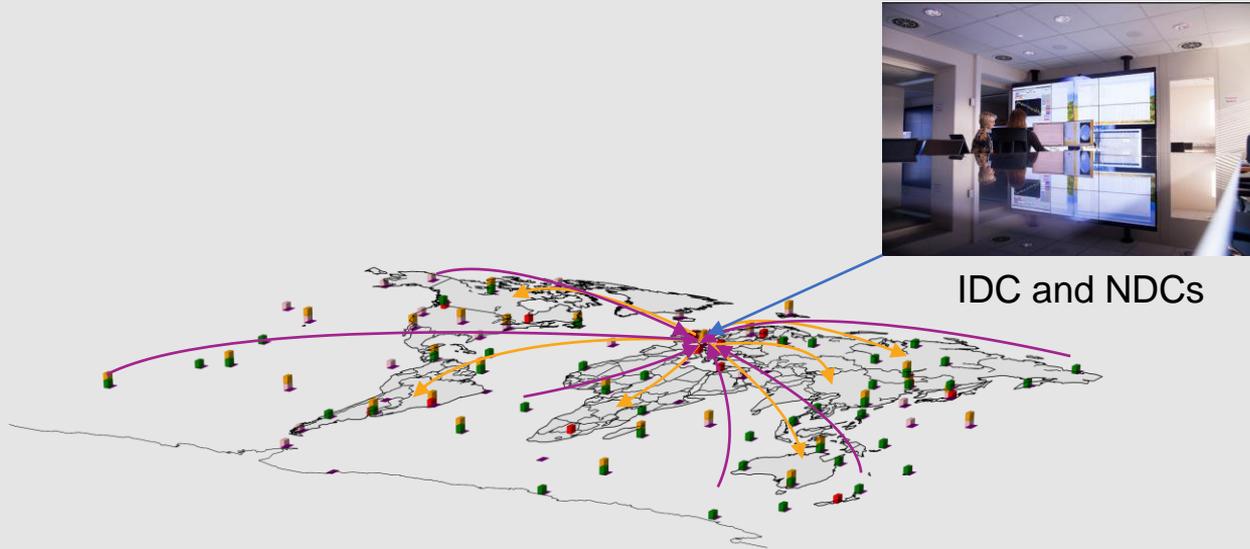
Re-measures archive filters or bottles
Generally more sensitive measurements
than at stations



RL03, Seibersdorf

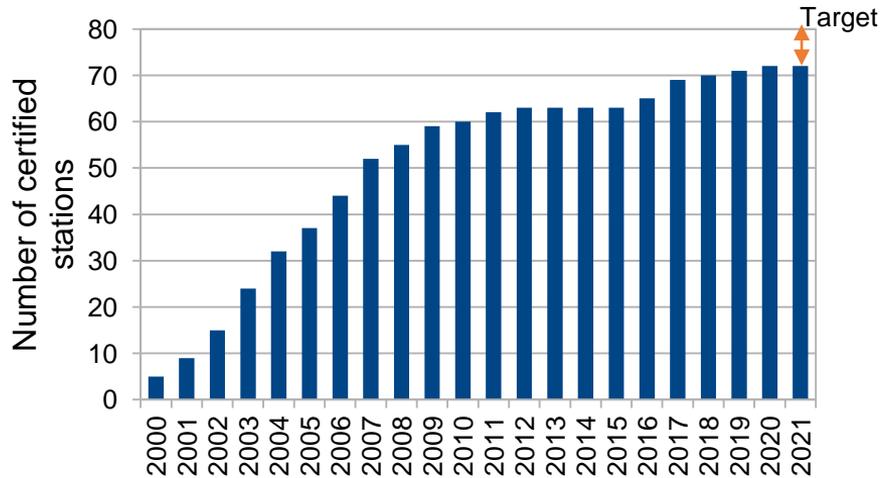


The IMS RN network 2021

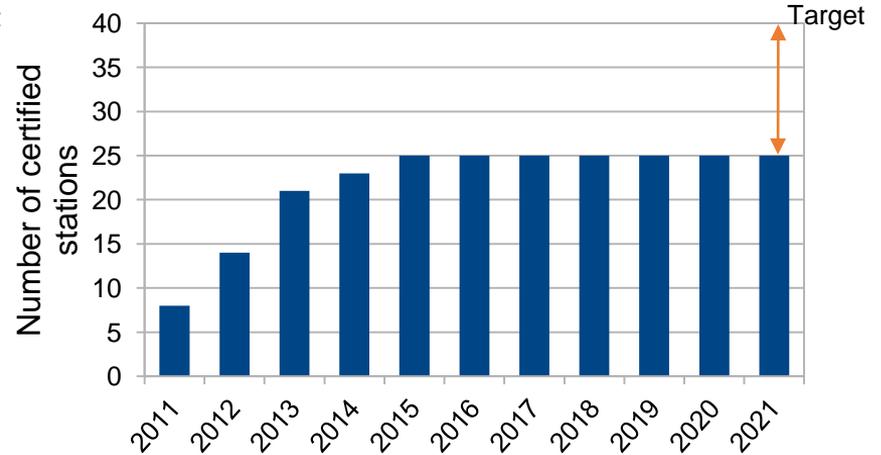


Station certification history

Particulates

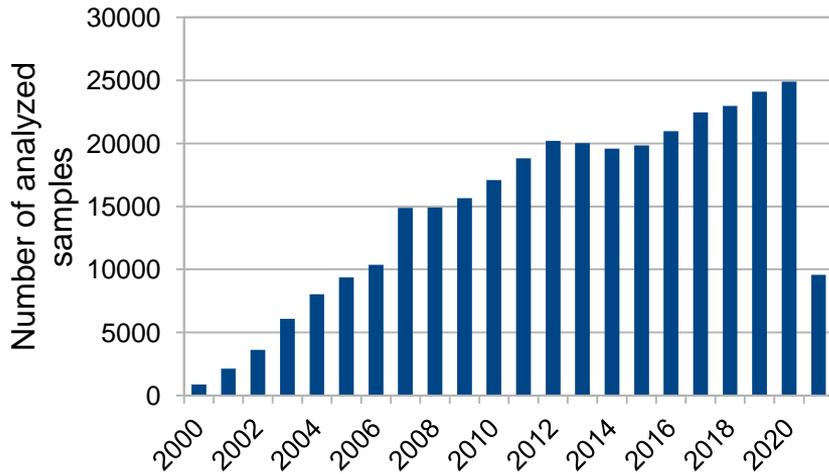


Noble gas

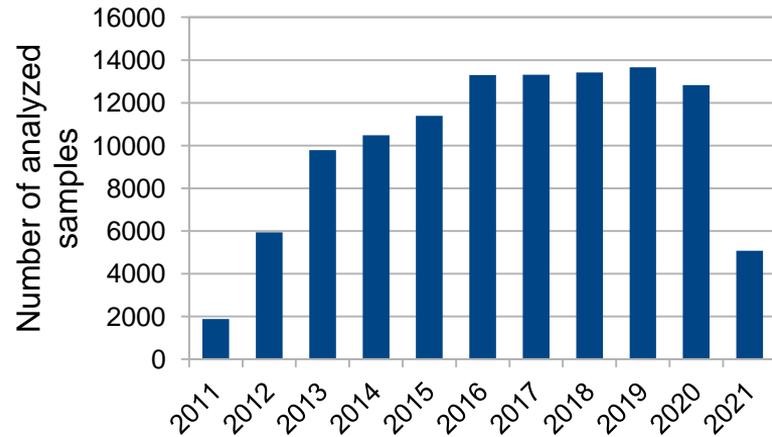


Sample analysis history

Particulate



Noble gas



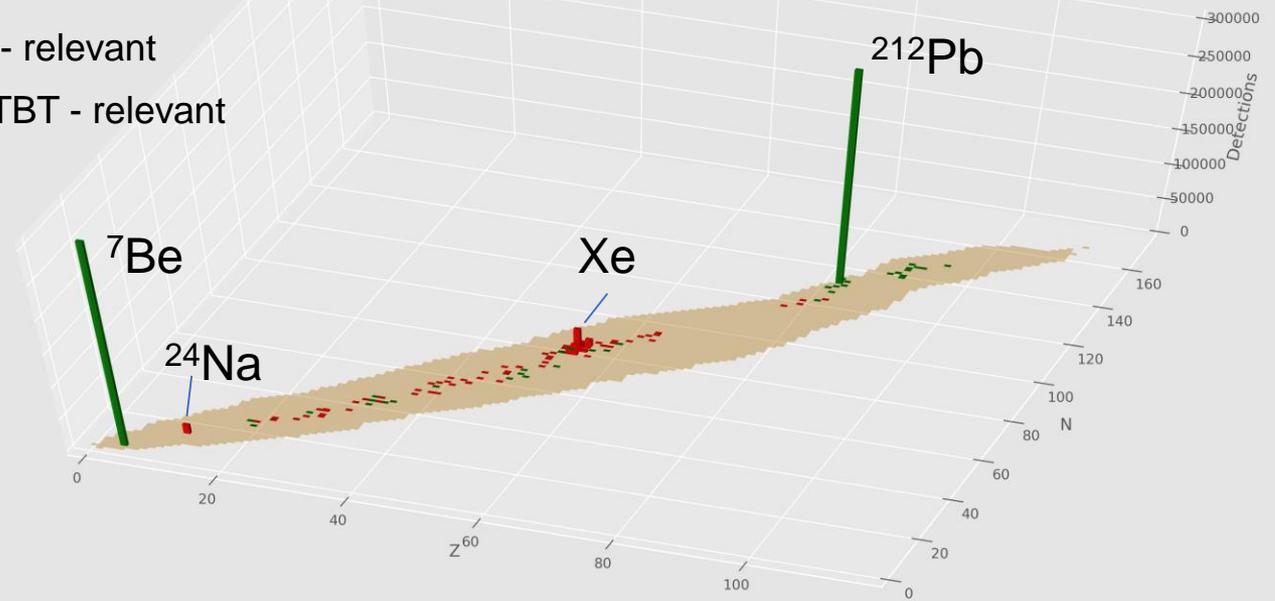


Examples of observations the last 20 years

Data extracted from the Swedish NDC, but using the results from the IDC reviewed analysis

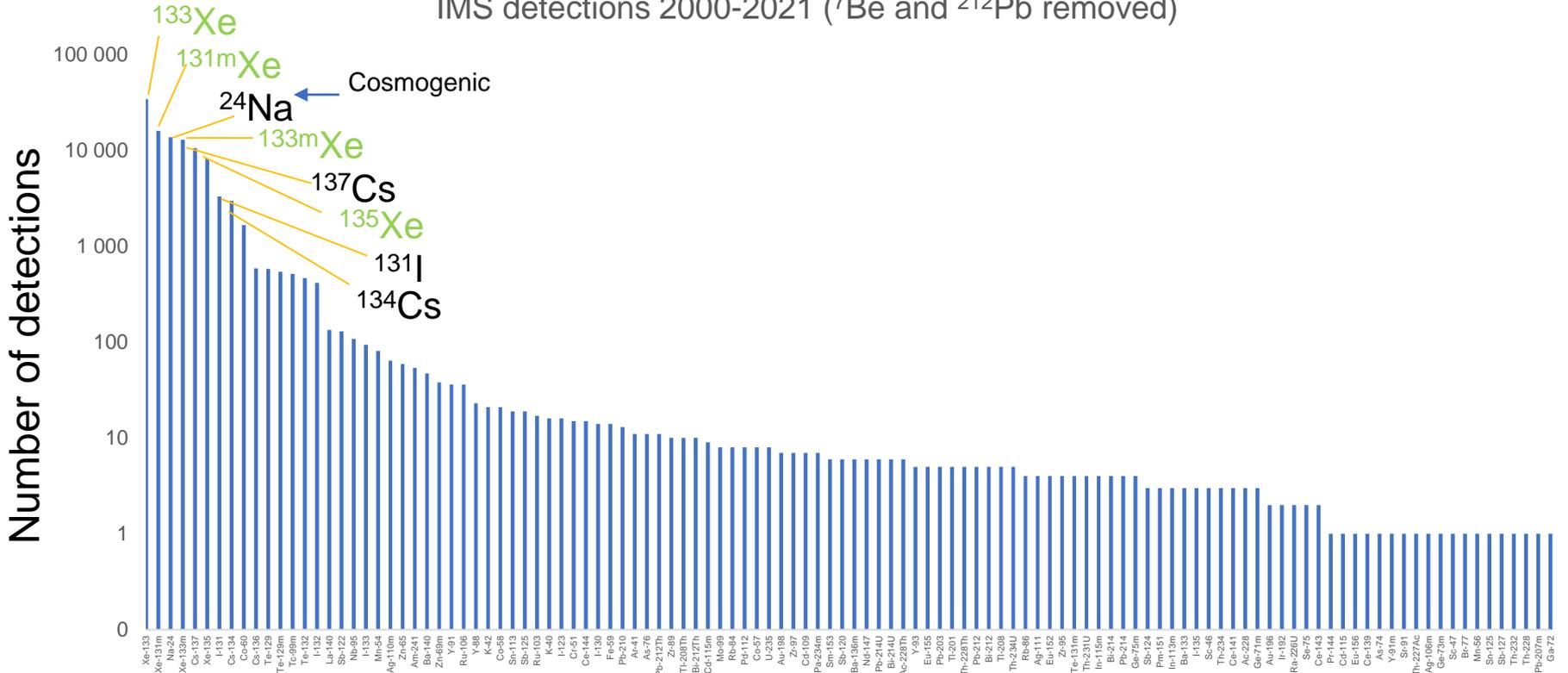
All reported detections 2000 - 2021

- CTBT - relevant
- Not CTBT - relevant



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IMS detections 2000-2021 (^7Be and ^{212}Pb removed)

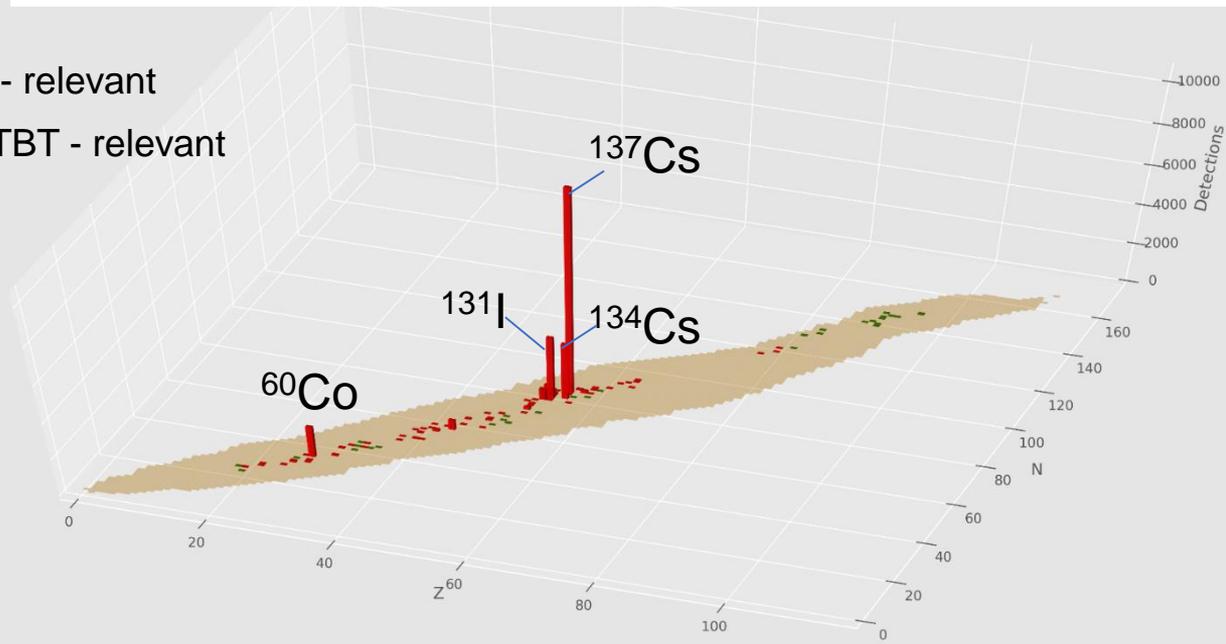


Disclaimer: The views expressed on this presentation are those of the author and do not necessarily reflect the view of the CTBTO

Detections 2000 – 2021, except ${}^7\text{Be}$, ${}^{212}\text{Pb}$, ${}^{24}\text{Na}$, and Xe

■ CTBT - relevant

■ Not CTBT - relevant

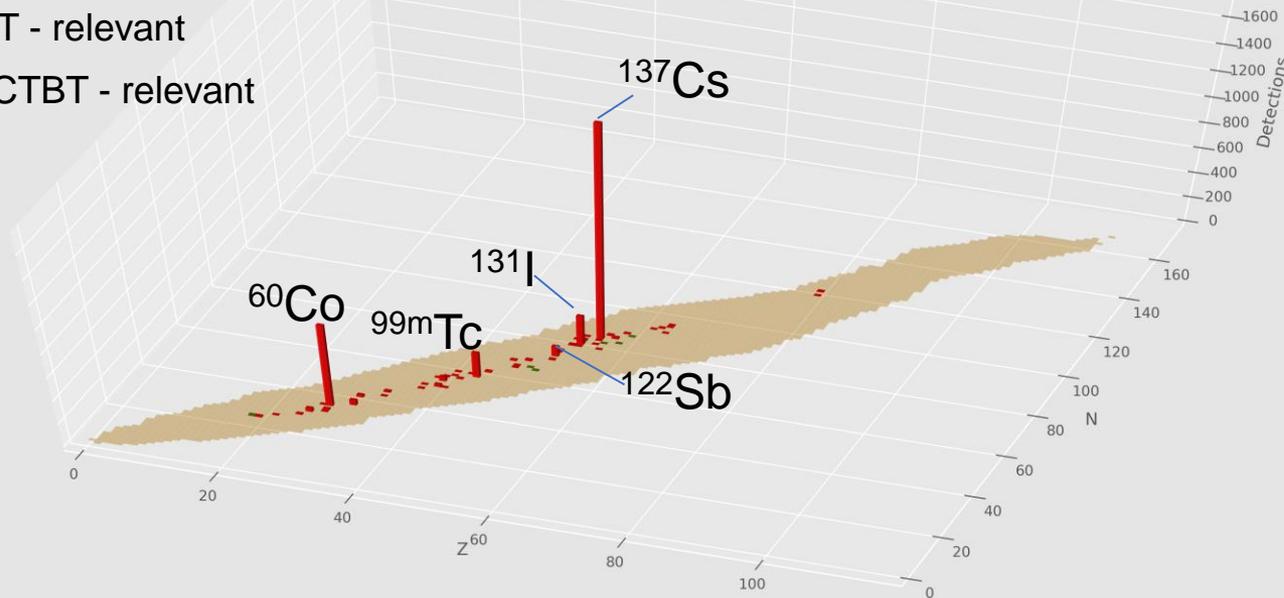


Disclaimer: The views expressed on this presentation are those of the author and do not necessarily reflect the view of the CTBTO

Detections 2000 – 2010, except ${}^7\text{Be}$, ${}^{212}\text{Pb}$, ${}^{24}\text{Na}$, and Xe

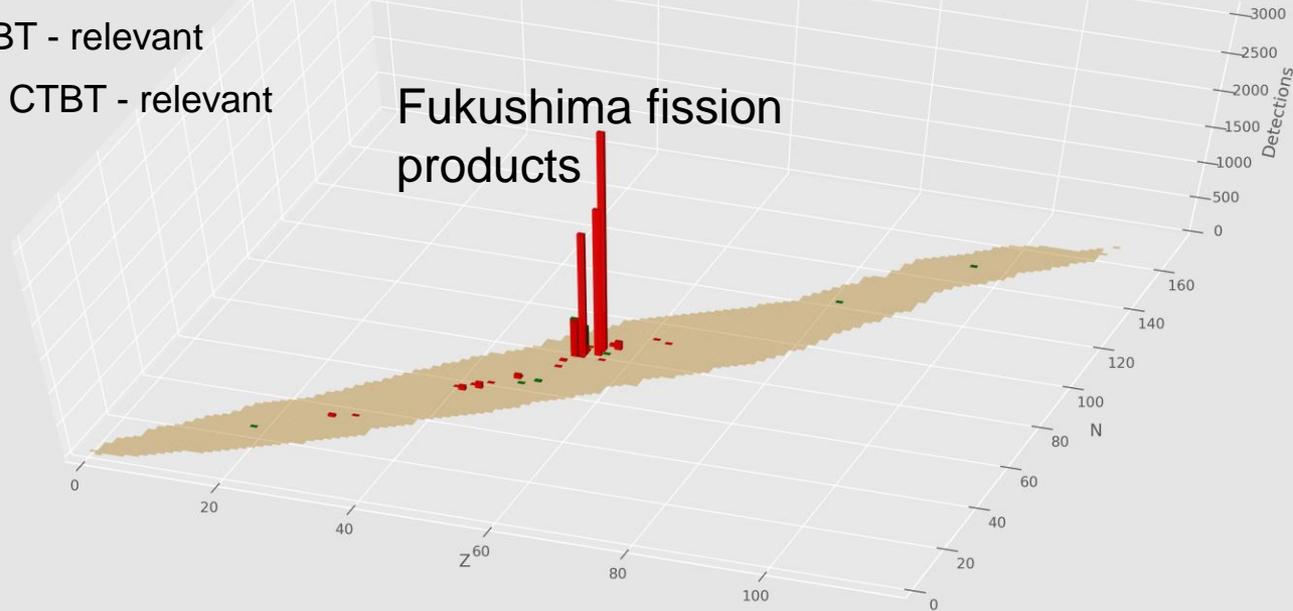
■ CTBT - relevant

■ Not CTBT - relevant



Detections 2011 – 2012, except ${}^7\text{Be}$, ${}^{212}\text{Pb}$, ${}^{24}\text{Na}$, and Xe

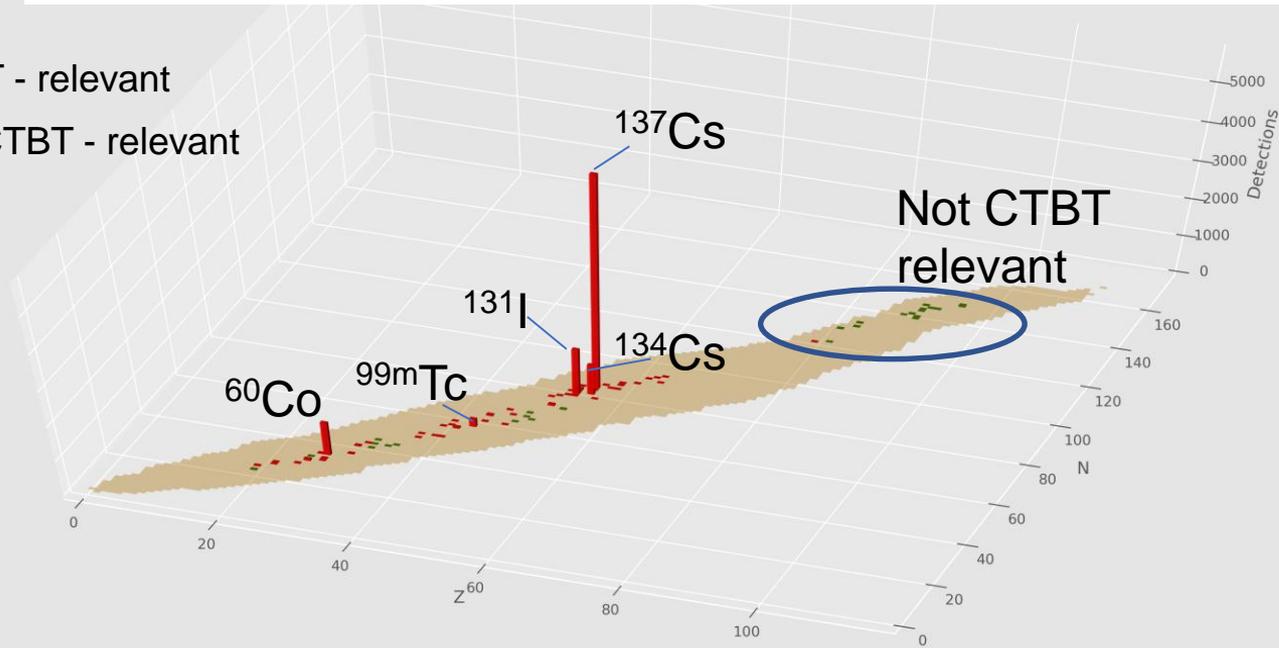
- CTBT - relevant
- Not CTBT - relevant



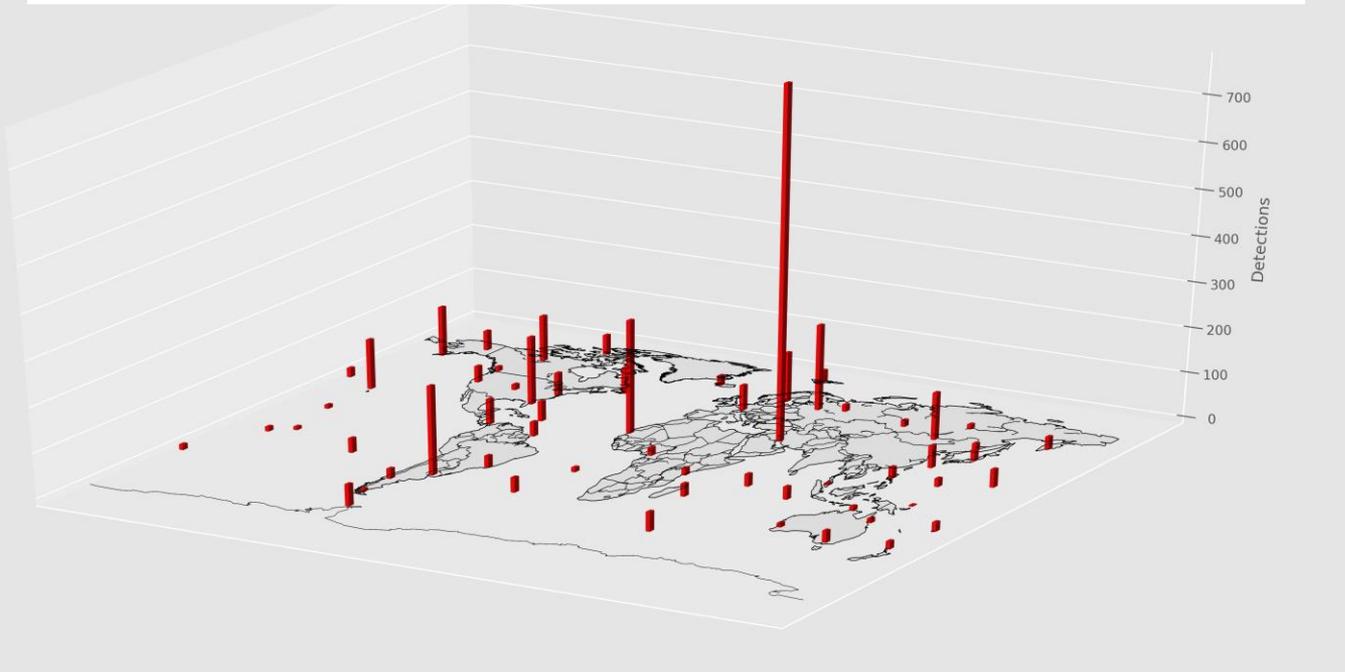
Detections 2013 – 2021, except ^7Be , ^{212}Pb , ^{24}Na , and Xe

■ CTBT - relevant

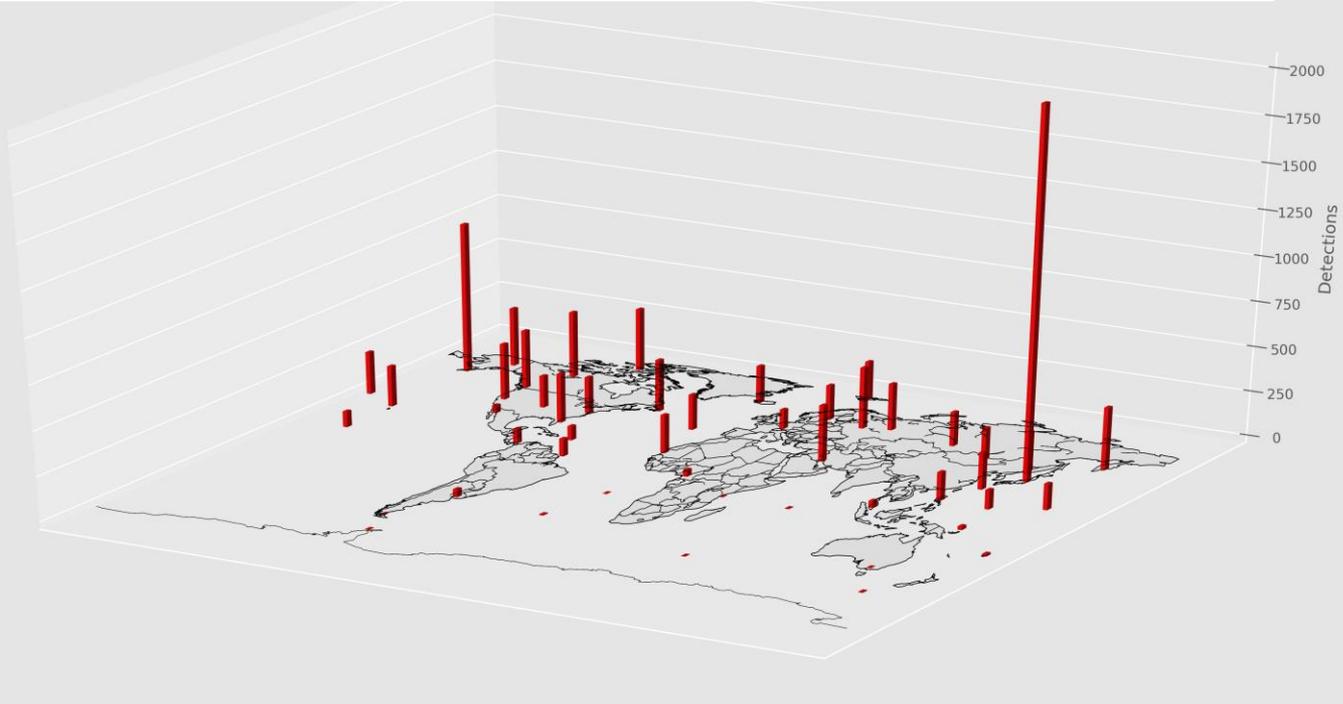
■ Not CTBT - relevant



Detections 2000 – 2010, except ^7Be , ^{212}Pb , ^{24}Na , and Xe

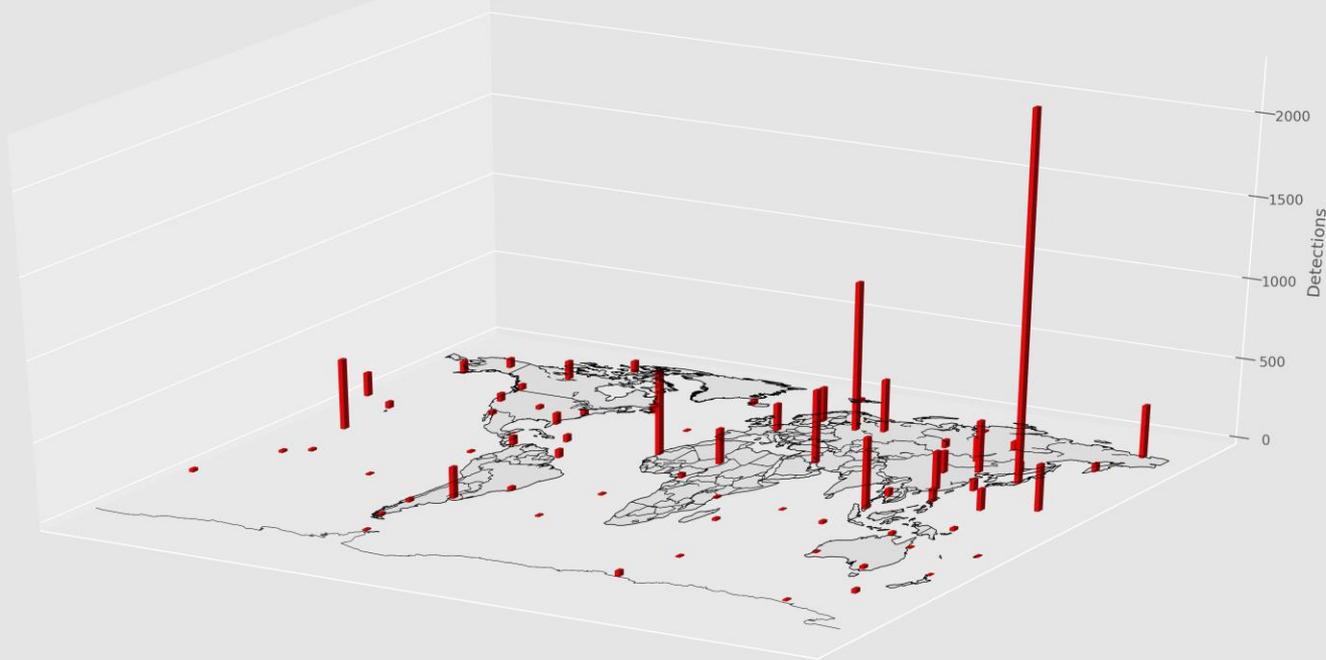


Detections 2011 – 2012, except ^7Be , ^{212}Pb , ^{24}Na , and Xe



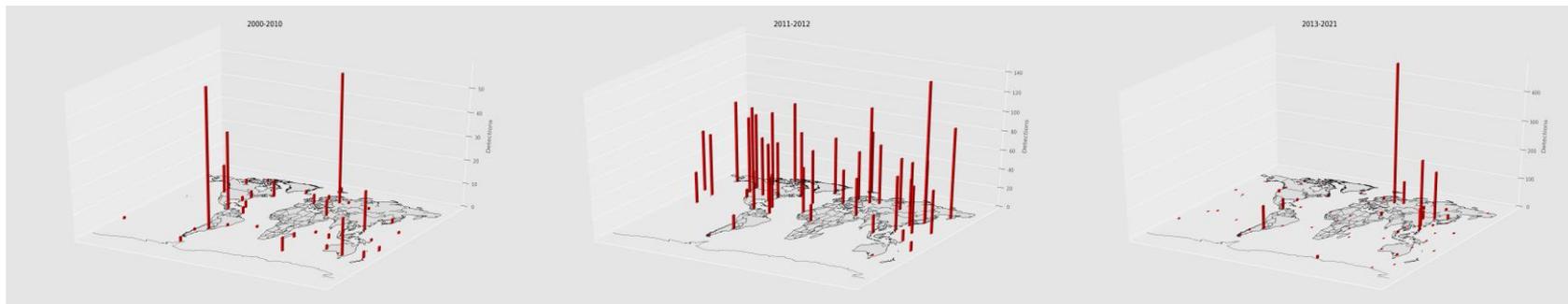
Disclaimer: The views expressed on this presentation are those of the author and do not necessarily reflect the view of the CTBTO

Detections 2013 – 2021, except ^7Be , ^{212}Pb , ^{24}Na , and Xe

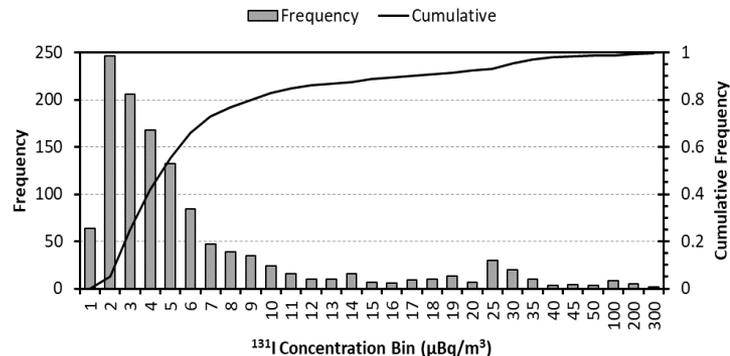


Disclaimer: The views expressed on this presentation are those of the author and do not necessarily reflect the view of the CTBTO

^{131}I – an important fission product with high background

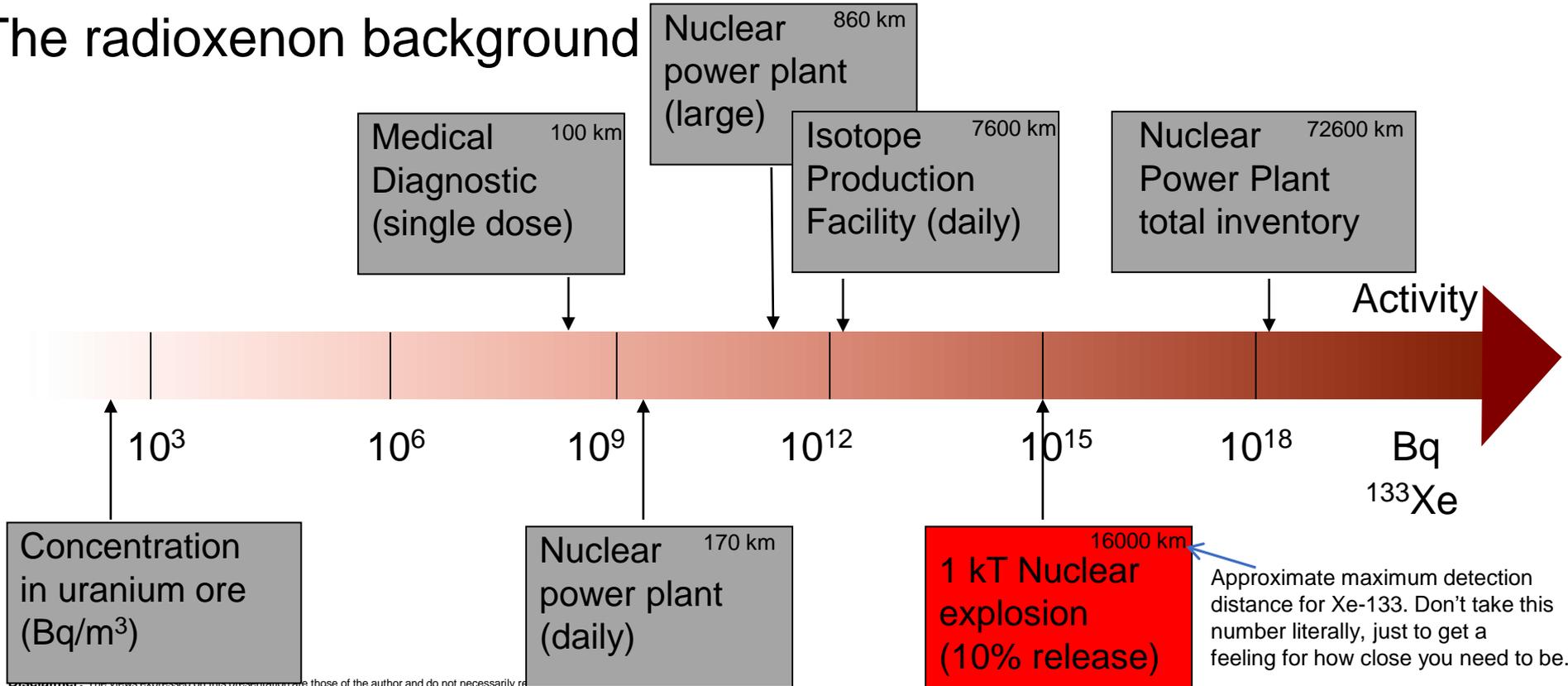


- Iodine the aerosol most likely to escape an UG test
- Observed frequently in IMS
- Many sources still unknown



Harry Miley, Paul Eslinger, Ramesh Sarathi, *Impact of environmental background on atmospheric monitoring of nuclear explosions*, talk at WOSMIP remote II, 2021

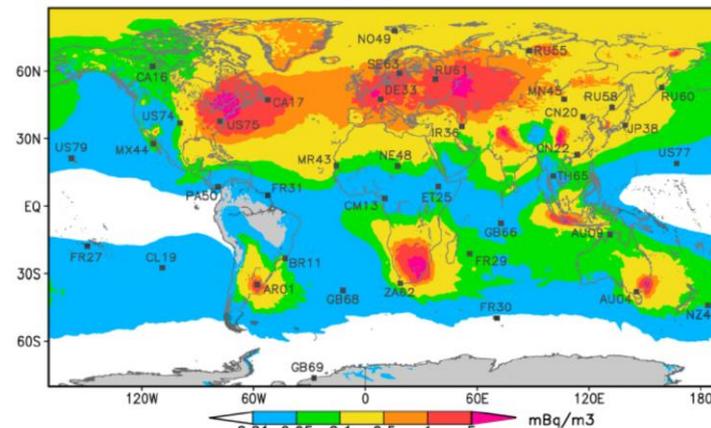
The radioxenon background



Disclaimer: The views expressed on this presentation are those of the author and do not necessarily reflect those of the Preparatory Commission of the CTBTO.

The radioxenon background

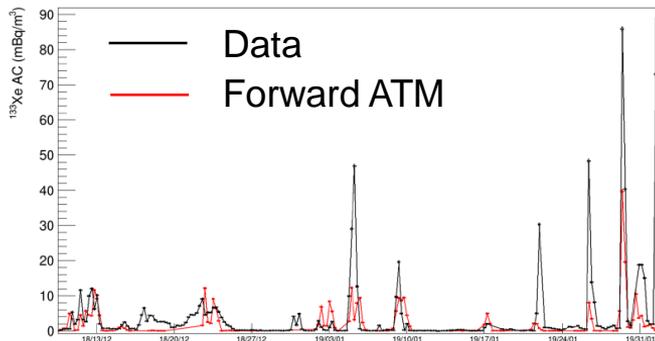
- The background is global, but variation between stations is high
- Many sources known, but not all
- The isotopic ratios can be close to NW ratios
- The background sources can mix with a true signal and disturb the ratios
- Pure ^{135}Xe observations
- Pure $^{131\text{m}}\text{Xe}$ observations



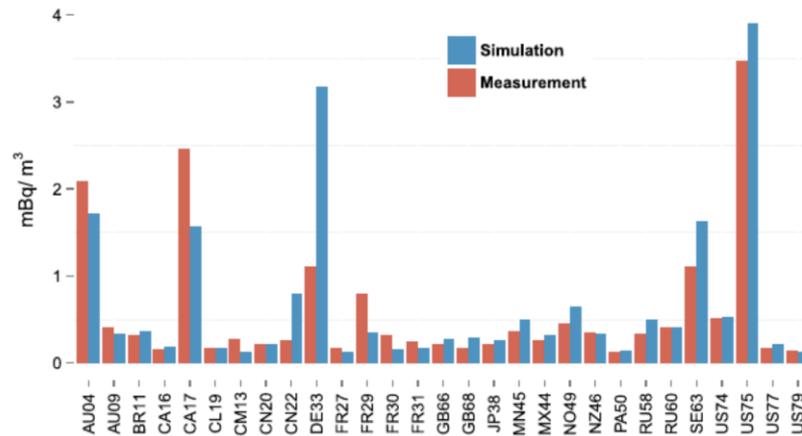
Map from : Achim, P., S. Generoso, M. Morin, P. Gross, G. Le Petit, and C. Moulin (2016), Characterization of Xe-133 global atmospheric background: Implications for the International Monitoring System of the Comprehensive Nuclear-Test-Ban Treaty, *J. Geophys. Res. Atmos.*, 121,4951–4966, doi:10.1002/2016JD024872

The radioxenon background

- The *average behaviour* of the radioxenon background can today be explained using known sources and ATM
- Individual cases more difficult



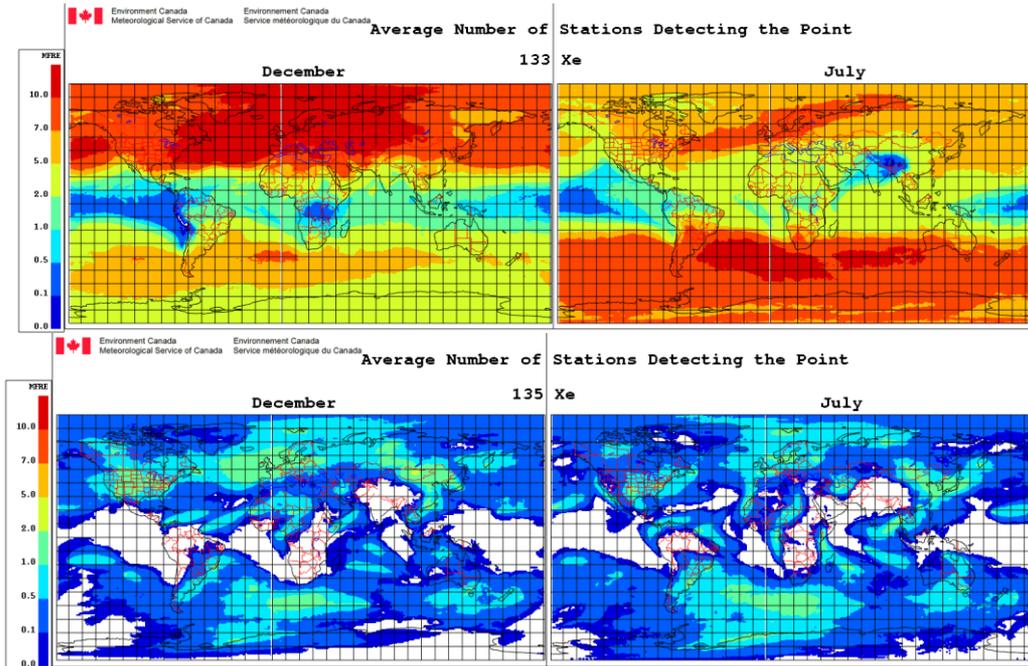
A. Ringbom et al., <https://doi.org/10.1007/s00024-020-02425-z>



Graph from: Achim, P., S. Generoso, M. Morin, P. Gross, G. Le Petit, and C. Moulin (2016), Characterization of Xe-133 global atmospheric background: Implications for the International Monitoring System of the Comprehensive Nuclear-Test-Ban Treaty, *J. Geophys. Res. Atmos.*, 121,4951–4966, doi:10.1002/2016JD024872

Network coverage for radioxenon

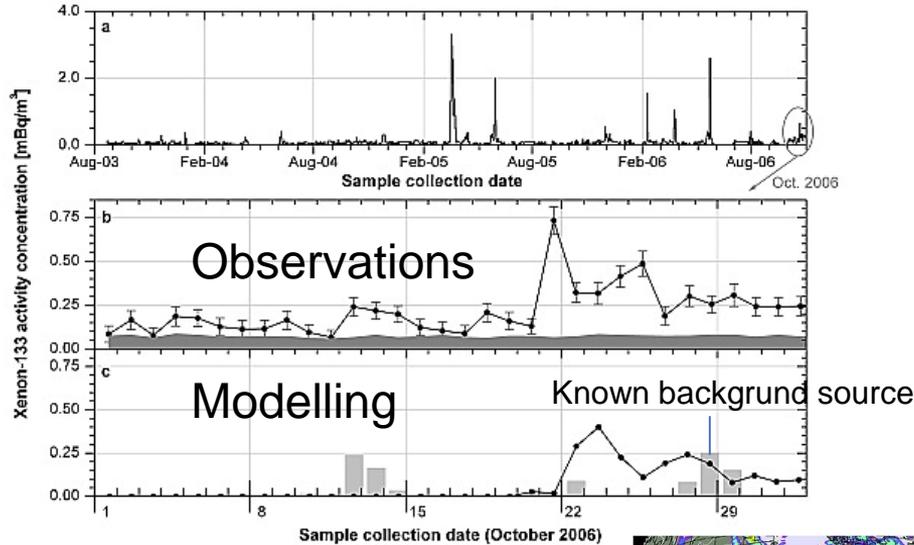
^{133}Xe
 $1.06 \times 10^{15} \text{ Bq}$



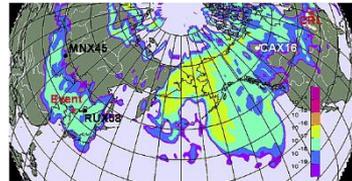
^{135}Xe
 $3.91 \times 10^{14} \text{ Bq}$

R. D'Amours and A. Ringbom, International Scientific Studies (ISS), Vienna, June 10-12, 2009

DPRK I - Yellowknife, Canada, October 2006

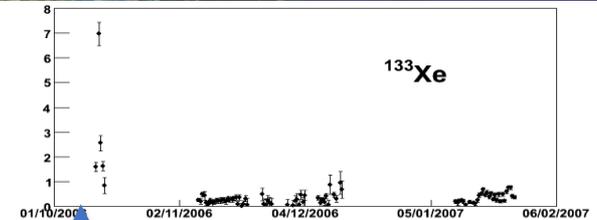
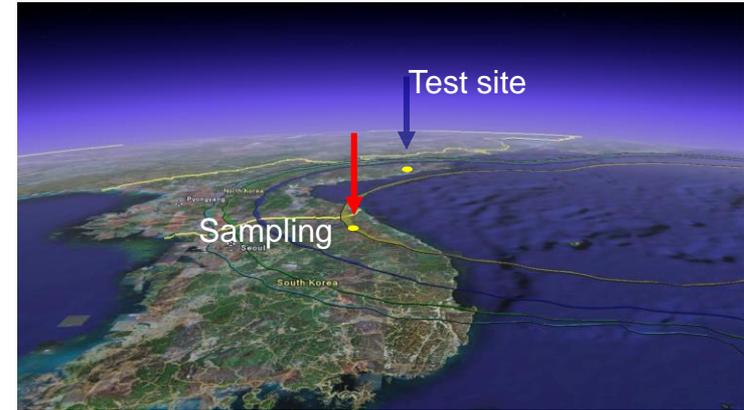


P. Saey et.al, *Geophysical Research Letters*, Volume: 34, Issue: 20, First published: 16 October 2007, DOI: (10.1029/2007GL030611)



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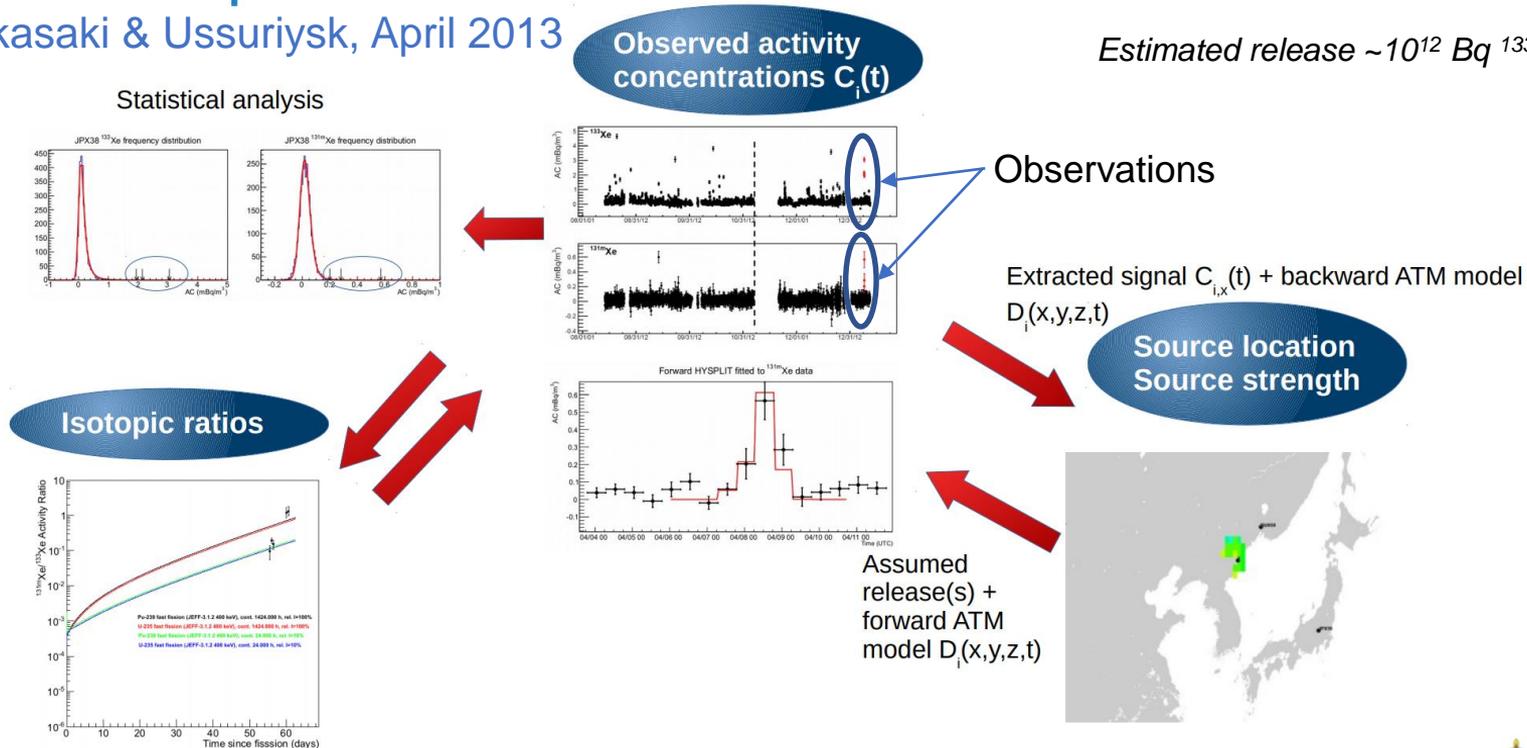
DPRK I – Measurement in ROK by Sweden and ROK



A. Ringbom et al., *Journal of Radioanalytical and Nuclear Chemistry* 282(3):773-779

DPRK III – Takasaki & Ussuriysk, April 2013

Estimated release $\sim 10^{12}$ Bq ^{133}Xe



**"Radioxenon detections in the CTBT international monitoring system likely related to the announced nuclear test in North Korea on February 12, 2013", Ringbom, et.al., <http://dx.doi.org/10.1016/j.jenvrad.2013.10.027>*

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Status of RN IMS – some conclusions

- The network is close to be complete, but important to get remaining radionuclide stations installed and certified
- Not all CTBT-relevant radionuclides are as relevant as others
- The background for particulates is relatively low, with a few exceptions.
- The radionuclide background is global and variable.
- Important to continue to identify background sources for xenon and iodine.
- The network xenon detection coverage needs to increase
- The released activity from a NT was overestimated when the network was designed. 33% of the DPRK tests detected.
- The methods used for location needs to be improved, including uncertainty estimates in ATM modelling.



Future development

The future development should be guided by a *network perspective*, based on a *scenario analysis* and using the *entire verification process**.

$$\text{Network verification power} = (D + L + R + T)/4$$

D - detection power

L - location power

R – rejection power

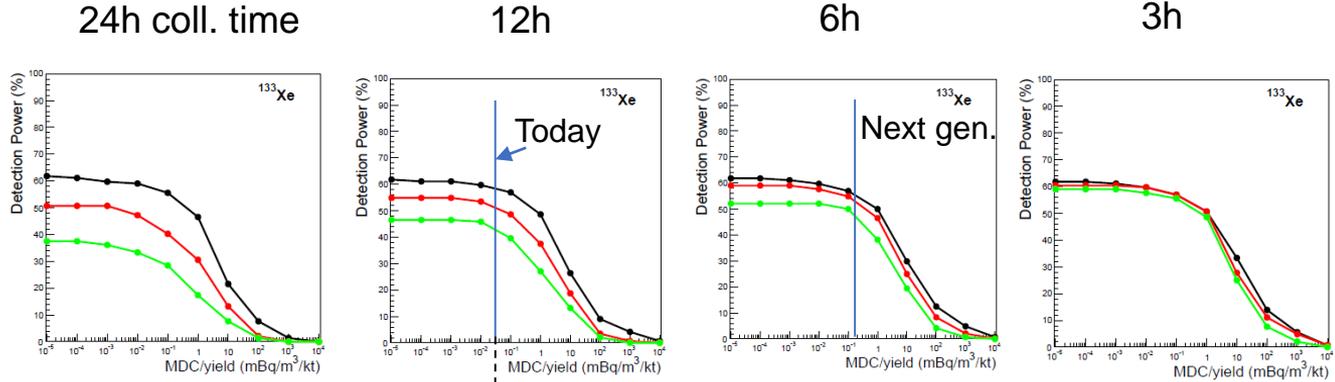
T – timing power

Example of result from an analysis using the network perspective:

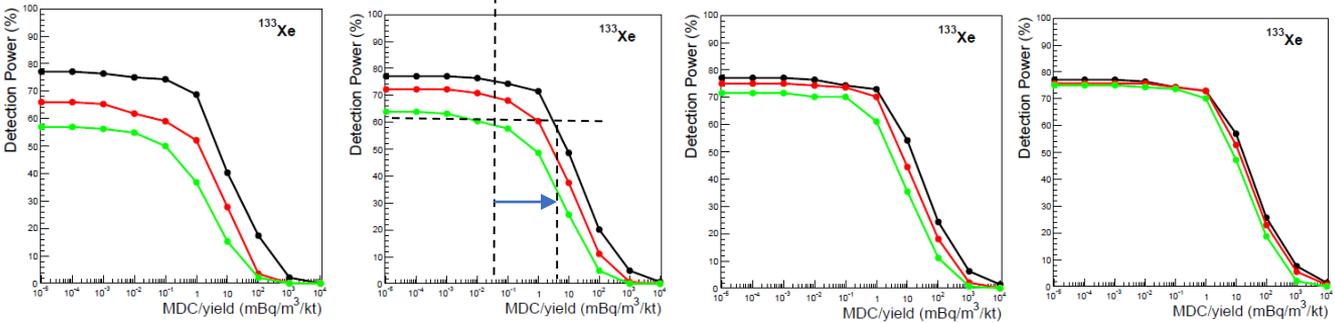
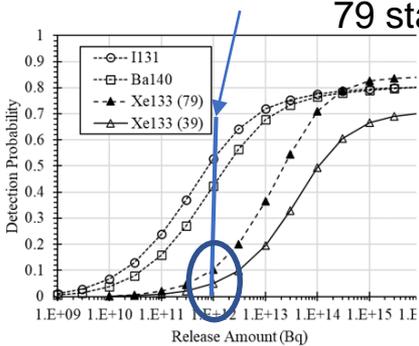
The same IMS xenon detection power for a 1 kT explosion is reached using 40 systems OR 80 systems with 30 times lower sensitivity.

FOI-R—3856—SE, *The impact of system characteristics on Noble Gas network verification capability.

Detection power* ^{133}Xe 39 stations



Xe levels observed from DPRK 79 stations



← Higher release or sensitivity

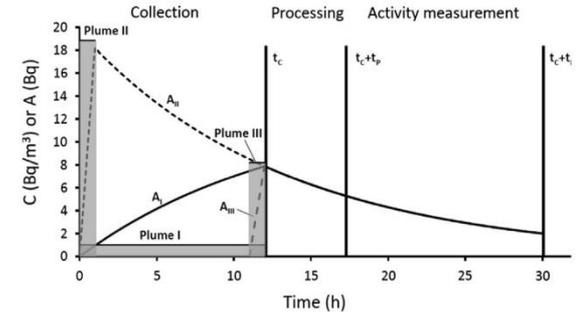
Harry Miley, Paul Eslinger, Ramesh Sarathi, talk at WOSMIP, 6/21/2021

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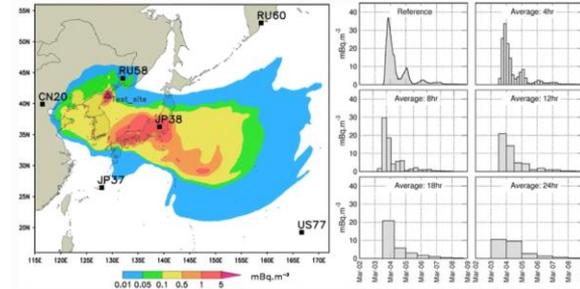
*FOI-R—3856—SE, *The impact of system characteristics on Noble Gas network verification capability.*

Impact of shorter collection time

- Increased number of samples
- Increased information on plume shape
- Better source location, but remains to be quantified.



A. Axelsson, A. Ringbom, Appl. Rad. Isot. 92 (2014) 12-17



Le Petit, G., Cagniant, A., Morelle, M. *et al.* Innovative concept for a major breakthrough in atmospheric radioactive xenon detection for nuclear explosion monitoring. *J Radioanal Nucl Chem* **298**, 1159–1169 (2013). <https://doi.org/10.1007/s10967-013-2525-8>

Next generation IMS NG systems



SAUNA III, Sweden

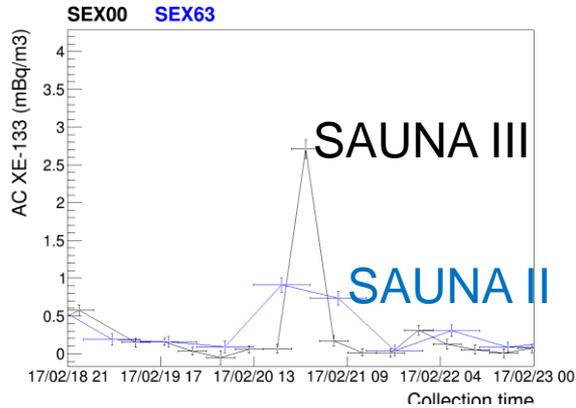


SPALAX NG, France

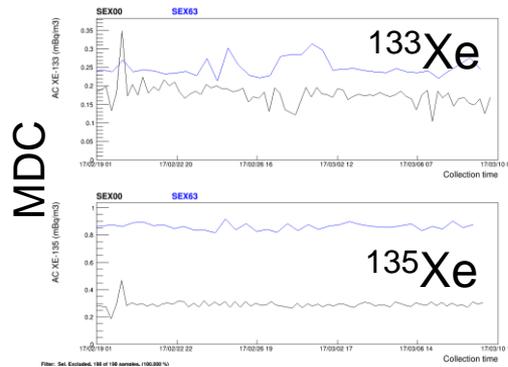


Xenon International, US

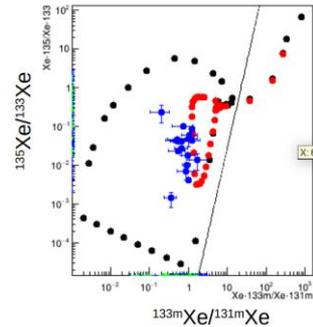
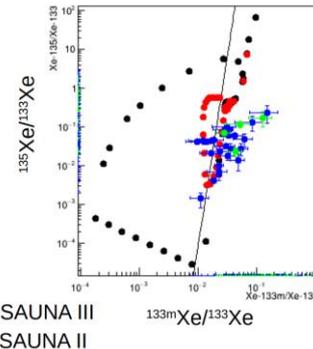
Next generation radioxenon systems



Increased time resolution



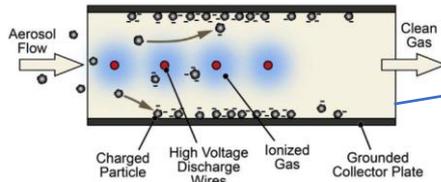
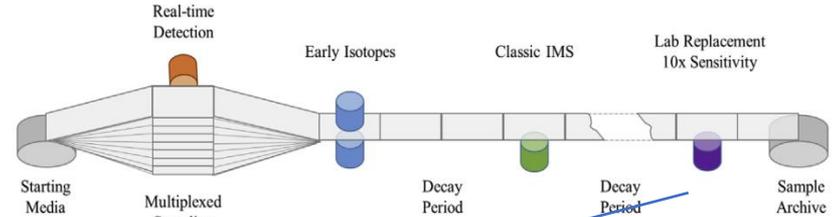
Improved sensitivity



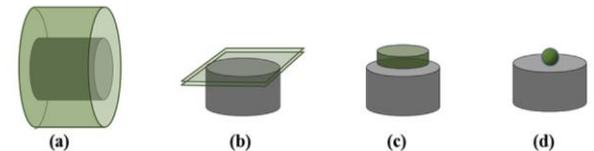
More samples

Development of particulate samplers

- Shorter sampling intervals
- Higher air flow
- Electrostatic precipitation (ESP)
- Dual gamma and/or gamma-gamma coincidence



$$MDC \approx C' \frac{e^{-\lambda_B t_d} \times \frac{(1 - e^{-\lambda_B t_c})}{\lambda_B}}{e^{-\lambda_A t_d} \times \frac{(1 - e^{-\lambda_A t_c})}{\lambda_A}} \times \frac{[^{212}\text{Pb}] \times FWHM}{V \times t_c \times \epsilon \delta \Omega}$$

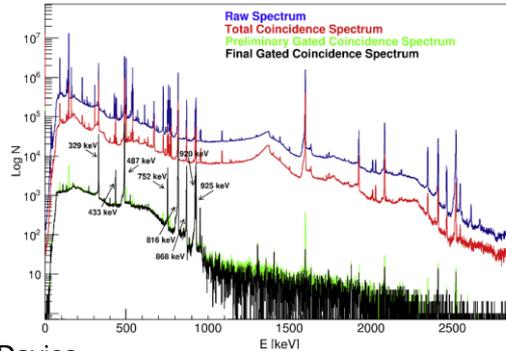
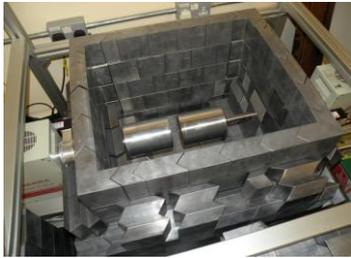


H. Miley et al., Design considerations for future radionuclide aerosol monitoring systems Journal of Environmental Radioactivity, Volumes 208–209, 2019, 106037, ISSN 0265-931X, <https://doi.org/10.1016/j.jenvrad.2019.106037>

Disclaimer: The views expressed on this presentation are those of the author and do not necessarily reflect the view of the CTBTO

Increased laboratory capabilities – coincidence detector systems

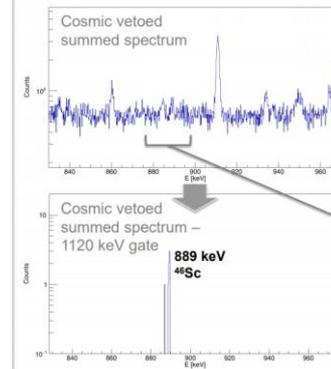
AWE



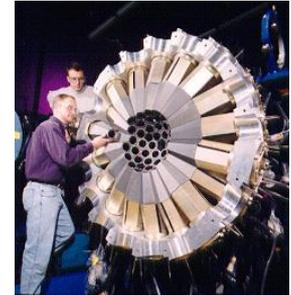
R. Britton, M. J. Jackson, A. V. Davies,
<http://dx.doi.org/10.1016/j.jenvrad.2015.07.025>
See also poster P3.1-303

- Improved sensitivity (veto, coincidence, and summing)
- Reduce biases from sample inhomogeneities
- Redundancy
- New data formats: list-mode

Health Canada,
Pawel Mekarski,
Dec 17, 2020



Gammaspere

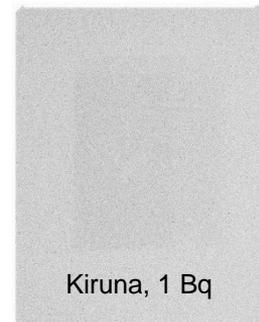
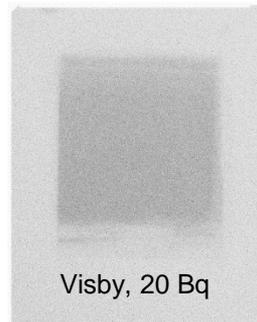
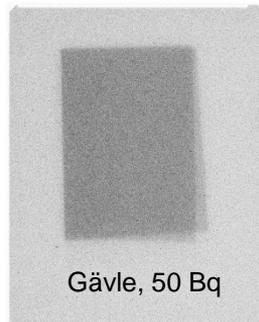


Argonne Physics Division - Low Energy Physics (anl.gov)



Increased laboratory capabilities – radiography

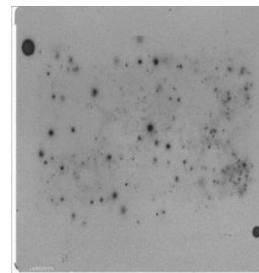
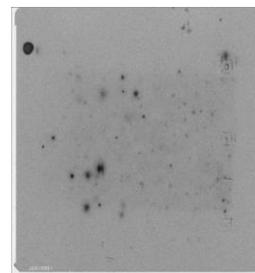
^{106}Ru , measured in Sweden 2017



- Information on production process
- Particle size distribution important for ATM
- Input how to interpret measured activity concentrations

Pictures provided by K. Ungar, Health Canada

IMS station in Takasaki, March 24 , 2011 $^{134,137}\text{Cs}$



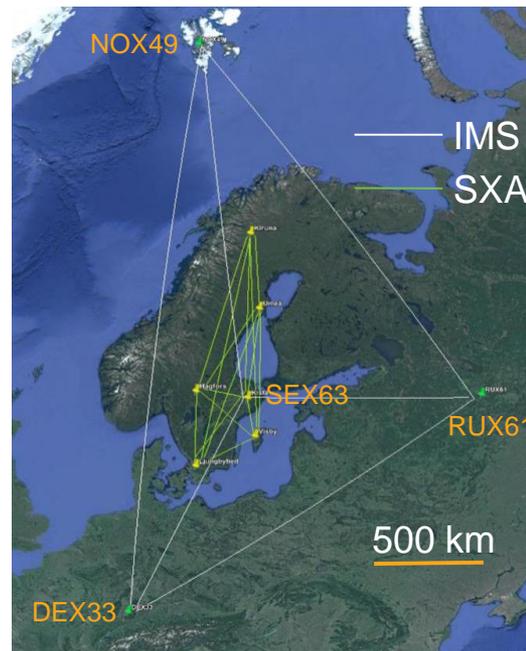
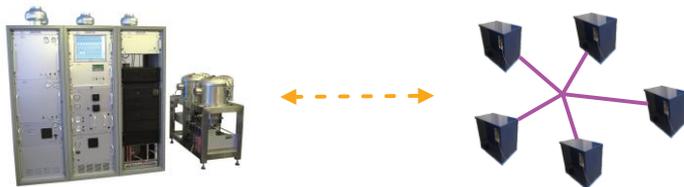
The SAUNA Q_B – array is the next step in remote sensing of activities involving nuclear fission that can improve..

Detection capability by decreasing average source-receptor distance and increasing coverage.

Location capability by increasing number of detecting sensors.

Categorization capability by increase the number of samples.

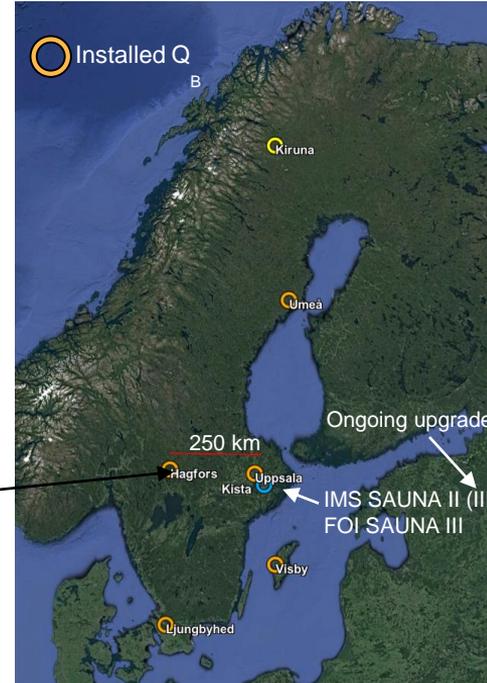
... at the same cost as a single state-of the art system like SAUNA III



The Swedish xenon array (units connected by green lines) shown together with nearby IMS radionuclide stations (white lines).



Five Q_B - array installed and running since May 5, 2021



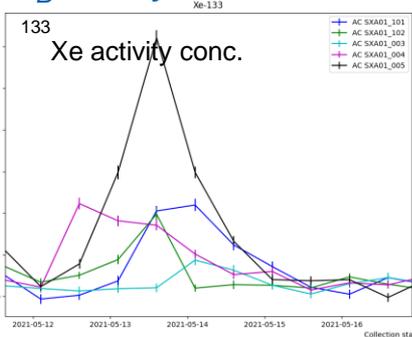
Installation of the first Q_B - unit in Hagfors, Sweden, in November 2020.

Example of the Q_B – array and SAUNA III detecting the same xenon plume

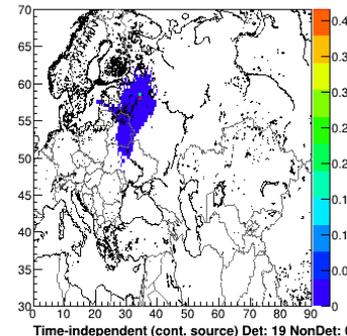
- A 2.5-day wide plume of ^{133}Xe hit Sweden in the middle of may 2021.
- All five Q_B - units in the array were affected, as well as the FOI SAUNA III in Stockholm.
- A Bayesian location analysis* was performed on the two data sets (19 Q_B – samples and 11 SAUNA III - samples), assuming a continuous source.
- The area of the resulting source location probability distribution is considerable smaller for the Q_B - array.



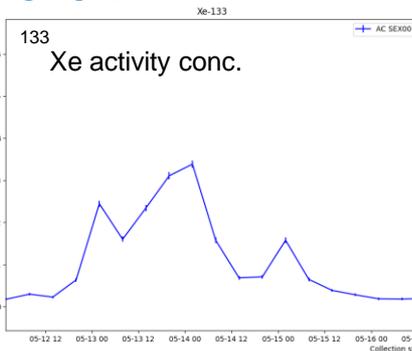
Q_B - array



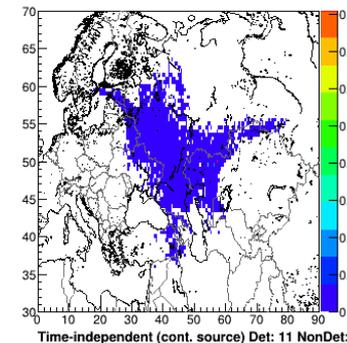
Continuous source Bayesian PDF



SAUNA III



Continuous source Bayesian PDF



*A. Ringbom and A. Axelsson, Poster at SnT2015, T1.3-P1.

Some thoughts for the future

- Development of equipment should take the entire verification process into account
- The next generation NG systems will improve categorization, detection and source term estimation
- Aerosol detection, including labs, is very sensitive, but there is still potential for improvement
- Upgrading IMS stations with array technology would be a major step
- Work can still be done on nuclear data and source term modelling
- Also true for network analysis, including ML and other techniques

